

The DLR-Institute of Space Systems (DLR-RY) and Fundamental Physics

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Institute of Space Systems Mission

.....The Institute develops concepts for innovative space missions on high national and international standard. Space based applications needed for scienctific, commercial and security-relevant purposes will be developed and carried out in cooperative projects with other research institutions and industry.....





Institute of Space Systems (DLR-RY) **Structure System Analysis** Space **Navigation Transportation System Analysis** and **Orbital Systems Control Systems Transport and System Test Propulsion** and **Systems** Verifiaction **Orbital Systems** and **Science Missions** Security **Avionic Systems Exploration Systems** Deutsches Zentrum für Luft- und Raumfahrt e.V. Folie 3 aG-WS, Pisa, 12.2.2010 in der Helmholtz-Gemeinschaft

AsteoridFinder / Start Phase B

Mission statement:

TheAsteroidFinder Mission observes the population of Near Earth Objects (NEOs), in particular IEOs (Inner Earth Objects / Interior to Earth's orbit objects) wrt:

- •Number of objects
- •Orbit distribution and orbit parameters (ephemeris)
- Scale distribution

- ✓ System Lead, Institute of Space Systems, DLR-RY
- Instrument development (telescope) and scientific lead, Institute of Planetary Research, DLR-PF
- 6 more DLR Institutes contributing: RM, RB, DFD, FA, ME, SC
- ✓ Scheduled launch: 2012

für Luft- und Raumfahrt e.V.

in der Helmholtz-Gemeinschaft

spacecraft

orbiting the Earth



▼ SSO (600 – 800 km), LTAN 06:00



AsteoridFinder Satellite

- ▼ 150 200 kg

- → 3 axes stabilised attitude
- ✓ no thrusters
- unregulated voltage: 18-24V
 power supply 285 W
- Kcommunication:
 S-band (omni-direktional) and X-band
- ✓ Passive thermal control system





Zodiacal and stray light



Folie 7

Instrument-Anforderungen und Möglichkeiten

- Images: 5 s⁻¹ (frames per second) (~200 ms exposure time)
- Stacking of up to 300 frames within 60 s by means of guide stars .
- → Requirement: $V_{lim} > 18,5$ mag:
 - ➤ No sunlight into the telescope
 - No remittance from earth (albedo) into the telescope
- ✓ CCD temperature (on focal plane): < -80 °C</p>
 - ➤ No sunlight on radiator
 - ➤ No earth infrared on radiator

> BUT:

Sufficient sunlight on solar panels for energy supply – needs optimum pointing to the sun





MASCOT (Marco Polo Surface Scout) Hayabusa 2

- - - ✓ Launch: 2014
 - Asteorid target: 1999JU3
 - ✓ Lander touch-down: 2018
 - Marco Polo NEO Sample Return Mission (ESA/Cosmic Vision - Launch: 2018/2019)
- ✓ French contribution: power and communication subsystems
- ✓ Lander able to adjust itself:
 - Guarantee of measuring position & mobility on the NEO surface (,hopping')
- Total mass of experimental equipment: 3 kg
 - \rightarrow 3 science instruments





Experimental devices

- 1. Wide angel camera (topography & geology of the landing terrains)
- 2. Mikroscope + IR spectrometer (mineralogy in µm-range)
- 3. Analyt. instrument (chemical composition) **or** Radiotomographer (innere structur)



System Analysis Space Transportation

Running projects

- Lead of Study "Bemannter Europäischer Raumtransport" BERT (Manned European Space Transportation)
- ELV-system design in cooperation with EADS Astrium and MT-Aerospace / nationale studies WOTAN and VENUS
- SpaceLiner (in 2009 EU-Project FAST20XX)
- Mikro-Launcher design and in cooperation with CNES within the Aldébaran Programme
- ✓ Systems analyses of propulsion engine, e.g. ceramic engine
- ✓ EU-Hypersonic projekts LAPCAT, ATLLAS ...

Facility:

Support for the Concurrent Engineering Facility







German Upper Stage Research Cooperation

- Research Cooperation Upper Stage
 - ✓ In cooperation with German launcher industry EADS- Astrium, MT Aerospace, University of Bremen ZARM and 4 DLR-Institutes

 - Focus: Re-ignitable cryogenic upper stage,
 5 Technology-Roadmaps:
 - ✓ Propellant Management Technology
 - ✓ Further Development of CFD-Tools
 - ➤ Simulation Feeding System
 - ➤ Composite Fibre Technology
 - ➤ Avionics Technology
- Testfacility in Bremen: Cryolab
 - ✓ Cryogenics LH2, LOX (50l), LN2 (1000l)
 - ✓ Vacuum chamber
 - ✓ Sloshing table











Re-entry Systems:







REX- Free Flyer

- Orbital system for re-entry and re-use
- Platform for experikments under weightlessness
- Highly variable experimental time (hours to weeks)
- Excellent µg-quality (~10⁻⁶g)
- Platform for the development of innovative reentry technology
- Re-enty control
- Landing control
- Variable size and weight
- Open for different launchers



DLR - Concurrent Engineering Facility

Concurrent Engineering (CE):



Sketch: ESA

Pre. S/W-Tool: Integrated Design Model (ESA)



Sketch: ESA

Power Thermal Structure Communication Data Mgmt. AOCS Propulsion

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Tools for AOCS/GNC Development



3-Axis Rotation Table



Magnetic Torquer Reaktion Wheels

Highly Precise Formation flight



- ➤ Motivation: Virtual structures (e.g. XEUS)
- - Sensors and actuators
- ➤ Development of test facilities





Autonomous Navigation for Exploration Missions



- ➤ Motivation:
 - ✓ Precise and soft landing
 - ➤ Autonomous orbit control
- - ✓ Terrain basede navigation





System Technology Central avionics

- ✓ Data management for space systems: satellites, launchers
- → Design: board computer and software
- ➤ Innovative concept: Network-Centric

STI Spacetec, Astrium IHP

Astrofein Technik (TET) Kayser Threde (TET) DLR: OS & SISTEC





Space Technology: Exploration Systems

- → Projects
 - ✓ ExoMars-Instrument HP³
 - Wheel development for ExoMars-Rover
- Studies
 - Cost-utilization-analysis of in-situ missios vs. sample return
 - ➤ Moon lander mission LAPIS

 - Contribution to Marco Polo Stuy (,Cosmic Vision') (ESA)
 - ✓ Marco Polo Lander ,MASCOT'
 - Contribution to ,Landing System' Study (ESA, with Alenia, Astrium)
- Contribution to HGF-Alliance "Planetary Development and Life"
- → Running mission contributions:
 - Mars Exploration Rover (NASA)







ExoMars-Rad

GG-WS, Pisa, 12.2.20

Marco Polo

Space Technology: Exploration Systems

Facilities

Centre for planetary landing Bremen

- ✓ Simulator LAMA for
 - Landing dynamics

 - Cooperation with

 - ✓ DLR-FA
- ✓ Planet simulation chamber







System Test and Verification

Location: Berlin-Adlershof

Facilities

- Simulation- and Test facilities for
 - ✓ Space environment, vacuum
 - ➡ Sun
- ✓ Mechanischal load
- ✓ EMC-Lab











Fundamental Physics Interests

- ➤ MICROSCOPE (Co-I)
- ✓ Inertial sensors
- ➤ S/C dynamics "science craft"
- ✓ S/C precision AOCS
- ✓ Precision thrusters (laser ablative thrusters)
- ✓ Quantum Optics in space
 - ✓ Quantum sensors / space atom interferometer (SAI)
 - ✓ Space clocks
 - → Optical links / time links
- ✓ Space Time Anisotropy
- ✓ Problems in deep space: communication, energy, propulsion



Challenge: S/C dynamics

- Study of appropriate attitude reconstruction method
 - ✓ Kalman filter, batch filter, ...
- Development of attitude estimation procedure
- Integration into Core Processes
- ✓ Physical attitude dynamics modelling
- ➤ Identification of satellite dynamics parameters





S/C Simulator Objectives

- Provide comprehensive simulation of the real system including science signal and error sources
- ✓ Provide simulation environment for control system performance validation
- Generate data needed to test data reduction methods
- ✓ Provide capability for identification of the satellite and instrument



Simulator Core Features

- Simulation of full satellite and test mass/experiment dynamics in six degrees of freedom by numerical integration of the equations of motion
- ✓ Multi-body system: e.g. STEP setup: satellite + 8 test masses → calculation of 117 states
- Consideration of linear and nonlinear coupling forces and torques between satellite and test masses/experiment as well as between test masses/experimental bodies
- ➤ Modelling of cross-coupling interaction
- Earth gravity model up to 360th degree and order, influence of Sun, Moon and planets can be included
- ✓ 5th order Runge-Kutta numerical integration, Bulirsch-Stoehr, Euler-Cauchy Several error sources are considered in the model:
- + misalignment and attitude errors
- + coupling biases
- + displacement errors



Force & Torque Modeling (1/2)

- Modeling of forces and torques acting on the satellite and test masses because of:

 - ✓ Control (forces and torques applied by the control system)
 - ✓ Interaction with the upper layers of the Earth atmosphere
 - ✓ Electromagnetic radiation
 - ➤ heat, radio communication emission
 - Absorption and reflection of radiation incident (Sun, Albedo, etc.)
 - ✓ Interaction with the magnetic field
 - ✓ Interaction (coupling) between satellite and experiment
 - ✓ From sensor and actuation systems
 - ✓ Gravitational coupling



Force & Torque Modeling (2/2)

➤ Modeling approaches:

- 1. Utilization **AND extension** of standard models
- 2. Derivation of parametric models from detailed FEM analysis of specific effects

✓ Standard models used:

- ✓ International Geomagnetic Reference Field (IGRF, IAGA)
- ➤ Mass Spectrometer Incoherent Scatter Model (MSIS, NRL)
 - Short-term variations of Earth atmospheric density (analysis of CHAMP mission data)



Verification of Simulator

- ➤ Comparison to analytical solution of simplified system
 - ✓ Simplified model renders ODE in Mathieu-Form
 - ✓ Verification by comparison of stability boundaries
- - ➤ Analytical description of uncoupled relative movement
- ✓ Verification of uncoupled attitude motion
- ✓ Verification with flight data (Gravity Probe B)



Simulator Architecture

- - ➤ Simulator for each mission is assembled from modules.
 - ➤ Modular design makes it easy to include new subsystems





MICROSCOPE

Micro-satellite à Trainée Componsée pour l'Observation du Principe d'Equivalence

Mission Parameter:

- Sun synchronous Orbit: 660 km
- Orbit-Excentricity : < 5 · 10⁻³
- Spin-Rate: variabel for modulation, der Orbit-Frequenz
- Signal frequency: $(\pi + 1/2) f_{orb}$ und $(\pi + 3/2) f_{orb}$
- Missions duration: 6 to 12 months
- Satellite mass: < 120 kg</p>
- CNES-Project (with contributions from DLR and ESA)



x-axis: sensitive axis z-axis: satllite spin axis

Problem: Clocks to explore space-time

- Redundant measurements
 - Measuring acceleration of S/C on geodesic via ranging and Doppler 7 tracking
 - Measuring redshift of clocks on-board S/C 7 for example Pioneer Anomaly

$$\frac{\Delta v}{v} = \frac{1}{c^2} \int_{20}^{90} \int_{AU}^{AU} a_{PA} \, \mathrm{d}x \approx 10^{-13}$$

Folie 31

Clock exploration does not depend on geodesic motion, independent from non-gravitational acceleration

- Cock exploration is cumulative 7
- Clocks automatically isolate the pure gravity sector
- Clocks represent an absolute DC-accelerometer



Problem: Time transfer (MWL)

- ➤ MWL (Microwave link to ISS) development from PRARE for time transfer between and ground
- ✓ Frequency comparison on the 10⁻¹⁶ level (230 fs per pass, 5 ps per orbit)
- ✓ 2 symmetric 1-way links carrying continous pseudo-noise coded signals
- ➤ High Ku-band chip rate (100 MChi) in order to increase resolution and suppotential multipath
- → Although not planned: ranging would be possible with $\lambda/1,000 = 24 \ \mu m$ accuracy.





Direct application

- → a path to relativisitc geodesy





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Future: Optical links

Optical transponders (on-board lasers, telescopes, timing receiver)

✓ Demonstrated over 0.17 AU (24 million km) with Messenger S/C and Mars Global Surveyor S/C (1-way)

➤ Nd:YAG laser, pulse rate 8 Hz

➤ Needs atmospheric correction: calibration can be done by ranging to near earth objects (e.g. LAGEOS) from differnt stations

	Messenger S/C	MOLA on Mars Global Surveyor S/C (1-way only)
range	2.4·10 ⁷ km	8·10 ⁷ km
pulsewidth	10 ns (up), 6 ns (down)	5 ns
pulse energy	16 mJ (up), 20 mJ (down)	150 mJ
repetition rate	240 Hz (up), 8 Hz (down)	56 Hz
laser power	3.84 W (up), 0.16 W (down)	8.4 W
beam divergence	60 µrad (up), 100 µrad (down)	50 µrad
receive area	0.042 m ² (up), 1.003 m ² (down)	0.196 m ²



Echo transponder for e.g. lunar laser ranging Time delay must be known



Asynchronous transponder for satellite laser ranging Repetition rate must be known

John J. Degnan, in Lasers, Clocks, and Drag Free...

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QUANTUS Collaboration





Quantum sensors

Based on: Ultra-precise optical metrology (1) $\delta v/v < 10^{-17}$ in the optical frequency domain Phase-sensitive atom interferometry (2) $\delta E/E < 10^{-19}$ due to the sub-microscopic quantum mechanical structure Needs experimental competence in: Processing ultra-cold atomic ensembles: (1) - "classical" laser cooled ensembles: ~ 1 μ K - Bose-Einstein Condensates (BEC): < 50 nK - ultra-cold molecules Measuring the phase highly precise: (2)- highly stable, phase-locked EM-oscialltors (RF-, THz-, optical) Calibrating oscillators: frequency comb Compa-Detector **Read-out**

2.2010

rator

Quantum Sensor

Long evolution time in μg

- BEC-TOF: 700 ms (meanwhile 850 ms)
- Thermal background disappeared
- Ca. 9,000 atoms
- Large extension: ca. 0.5 mm





Outlook (Experiments)

- Test of the Equivalence Principle with quantum objects
 - Freely falling quantum probes / distinct atomic species
 - CAPRICE experiment within QUEST-Programme (Quantum Engineering and Space-Time Research)
 - *M. Kasevich* et al. (2007):
 - Atom interferometer height: ca. 10 m
 - Wavepackage separation: > 10 cm
 - statistical accuracy:
 - systematic uncertainty:

 $\delta g/g < 10^{-15}$ $\delta g/g < 10^{-16}$



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Summary: DLR-RY competences

- ✓ System apporach: AsteoridFinder: DLR compact satellit "Sciencecraft"
- Cooperation with other DLR space related institutes for planetary research, GSOC/operations, robotics /mechatronics, communication / navigation, technical physics, light structure mechanics, construction, and S/W development
- ✓ System competence at RY:

 - ➤ AOCS / precision attitude control (NR)
 - ➤ S/C dynamic simulation
 - avionics

 - Mission analysis / concurrent engineering (OR / SA)



