

S3 - Lezioni 1 e 3 Marzo 2016:

- Confronto tra gli effetti della forza e del momento della forza; dimostrazione che se la forza è centrale il moto deve svolgersi in un piano
- Calcolo della energia potenziale gravitazionale a grande distanza e vicino alla superficie della Terra. Importanza del segno nei due casi.
- Richiami sullo sviluppo in serie di potenze
- Forze dissipative e loro effetto
- Forza di resistenza dell'aria su un satellite in orbita e su un corpo in caduta
- Forza di Archimede

$$\vec{F} = m\vec{a} = m \frac{d\vec{v}}{dt} = \frac{d}{dt}(m\vec{v})$$

$\vec{p} = m\vec{v}$

$\vec{L} = \vec{r} \times \vec{p}$

$\vec{N} = \frac{d\vec{L}}{dt}$

$\vec{N} = \vec{r} \times \vec{F}$

\vec{p} quantità di moto lineare

\vec{L} quantità di moto angolare (momento angolare)

MOTO PIANO

Forza centrale $\Rightarrow \vec{N} = 0 \Rightarrow \frac{d\vec{L}}{dt} = 0 \Rightarrow \vec{L} = \text{costante}$

Forza conservativa \rightarrow Energia potenziale

$U(r) = -\frac{GMm}{r}$

$U(P) = U(R+z) = -\frac{GMm}{R+z}$

$z \ll R$

$\epsilon \ll 1$

$\epsilon = 10^{-1}$
 $\epsilon^2 = 10^{-2}$
 $\epsilon^3 = 10^{-3}$

$\frac{z}{R} \ll 1$

$f(x)$ ($x \ll 1$) **parametro di piccolezza** SVILUPPI IN SERIE DI POTENZE $f = (1+\epsilon)^{-1}$

$f(x) \approx f(0) + x \frac{df}{dx} \Big|_{x=0} + \frac{x^2}{2} \frac{d^2f}{dx^2} \Big|_{x=0} + \frac{x^3}{3!} \frac{d^3f}{dx^3} \Big|_{x=0} + \dots + \frac{x^n}{n!} \frac{d^n f}{dx^n} \Big|_{x=0} + O(x^{n+1})$

$n! = n(n-1)(n-2) \dots (n-(n-1))$

$U(P) = -\frac{GMm}{(R+z)} = -\frac{GMm}{R(1+\frac{z}{R})} = -\frac{GMm}{R} \frac{1}{1+\epsilon} = -\frac{GMm}{R} (1+\epsilon)^{-1}$

$\epsilon = \frac{z}{R} \ll 1$

$f(\epsilon) = (1+\epsilon)^{-1} \approx f(0) + \epsilon \frac{df}{d\epsilon} \Big|_{\epsilon=0} + \frac{\epsilon^2}{2} \frac{d^2f}{d\epsilon^2} \Big|_{\epsilon=0} + O(\epsilon^3) \approx 1 - \epsilon + \frac{\epsilon^2}{2} + O(\epsilon^3)$

$\frac{d^3(1+\epsilon)^{-1}}{d\epsilon^3} = -1(1+\epsilon)^{-2} \cdot \frac{d(1+\epsilon)}{d\epsilon} = -(1+\epsilon)^{-2} \cdot 1 = -(1+\epsilon)^{-2}$
 $\frac{d^2(1+\epsilon)^{-1}}{d\epsilon^2} = + (1+\epsilon)^{-3} \cdot \frac{d(1+\epsilon)}{d\epsilon} = + (1+\epsilon)^{-3}$

energia potenziale di m sulla sup. della Terra

$U(P) = -\frac{GMm}{R} \left(1 - \frac{z}{R} + \frac{1}{2} \frac{z^2}{R^2} + O\left(\frac{z^3}{R^3}\right) \right) \approx -\frac{GMm}{R} \left(1 - \frac{z}{R} + \frac{1}{2} \frac{z^2}{R^2} \right) - U(\infty)$

$U(R+z) = -\frac{GMm}{R} \approx -U(R) = -\left(\frac{GMm}{R}\right)$

$U(R+z) + \frac{GMm}{R} \approx + \frac{GMm}{R^2} z - \frac{GMm}{2R^3} z^2$

$U(R+z) - U(R) \approx mgz - mgz \left(\frac{z}{2R}\right) \approx mgz \left(1 - \frac{z}{2R}\right) \approx \frac{2}{64} \times 10^{-5} \approx 3 \times 10^{-6}$

$U(z) = mgz > 0$

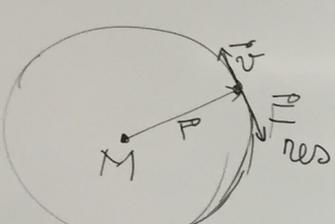
$g = \frac{GM}{R^2} \approx \frac{3.99 \times 10^{14}}{(6.4)^2 \times 10^{12}} \approx \frac{3.99}{(6.4)^2} \times 10^2 \text{ ms}^{-2} \approx \underline{\underline{9.8 \text{ ms}^{-2}}}$

$(GM)_{\oplus} = 3.99 \times 10^{14} \text{ m}^3 \text{ s}^{-2} \text{ kg}^{-1} \text{ K}_J$

Es: $\frac{40 \text{ m}}{2.64 \times 10^6} \approx \frac{2}{64} \times 10^{-5} \approx 3 \times 10^{-6}$

0.000003

$U(z) = mgz$

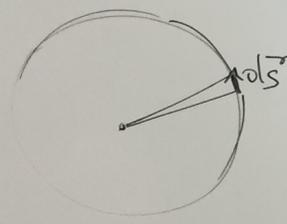


$v = \omega r = \sqrt{\frac{GM}{r}}$

$\omega^2 r^3 = GM$

$P = \frac{2\pi}{\omega}$

$E = -\frac{GMm}{2r}$



$dL = \vec{F}_{res} \cdot d\vec{s} = \vec{F}_{res} \cdot \vec{v} dt$

$\frac{dL}{dt} = \vec{F}_{res} \cdot \vec{v} = -F_{res} v$

energie

potenza

$\frac{dE_{grav}}{dt} = \frac{dL}{dt} = -F_{res} v$

$\frac{d}{dt} \left(-\frac{GMm}{2r} \right) \leftarrow \frac{dE_{grav}}{dt} = \frac{dL}{dt} = -F_{res} v$

FORZE DISSIPATIVE E LORO EFFETTO

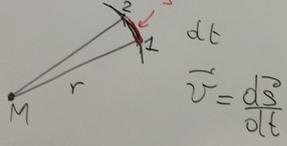


$P = \frac{2\pi}{\omega}$ $\omega^2 r^3 = GM$

$v = \omega r = \sqrt{\frac{GM}{r}}$

$E(r) = \frac{GMm}{2r}$ $\frac{dE_{grav}}{dt} ?$

$dL = \vec{F}_{res} \cdot d\vec{s} = F_{res} \cdot \vec{v} dt$



$\frac{dL}{dt} = -F_{res} v = -F_{res} \omega r$

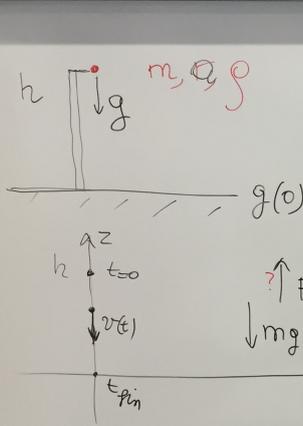
$\dot{r} = -\frac{2}{\omega} \left(\frac{F_{res}}{m} \right) < 0$ $\dot{r} = -2 F_{res} \omega r^3 = -\frac{2 F_{res} \omega r^3}{GMm}$

$E(r) = -\frac{GMm}{2r} = \left(-\frac{GMm}{2}\right) r^{-1}$

$\frac{dE_{grav}}{dt} = \left(-\frac{GMm}{2}\right) (-r^{-2} \frac{dr}{dt})$

$\frac{dE_{grav}}{dt} = \frac{GMm}{2r^2} \left(\frac{dr}{dt}\right)$

$\dot{E}_{grav} = -F_{res} \omega r$



$m = \left(\frac{4}{3}\pi a^3\right) \rho$ $v(0) = 0$

$F_{res} = \left(\frac{1}{4}\right) \rho_{aria} (\pi a^2) v^2$

$E(t=0) = E(t_{fin})$ $\dot{U} = mgh = \frac{1}{2} m v_{fin}^2$

$v(t_{fin}) = \sqrt{2gh}$

$F_{res} \ll mg$

$1 \Rightarrow \frac{(F_{res})_{max}}{mg} \approx \frac{1}{2} \frac{\rho_{aria} \pi a^2 g h}{\frac{4}{3} \pi a^3 \rho} = \frac{3}{8} \frac{\rho_{aria}}{\rho} \frac{h}{a} \approx 0.1$

$(F_{res})_{max} = \frac{1}{4} \rho_{aria} \pi a^2 v(t_{fin})^2 = \frac{1}{4} \rho_{aria} \pi a^2 g h$

$h = 15 \text{ m}$
 $\rho_{aria} \approx 1.2 \text{ kg/m}^3$
 $\rho = 2.7 \text{ g/cm}^3$
 $a = 2.5 \text{ cm}$

$F_A = g \left(\frac{4}{3}\pi a^3\right) \rho_{aria}$

$\frac{F_A}{mg} = \frac{g \frac{4}{3}\pi a^3 \rho_{aria}}{g \frac{4}{3}\pi a^3 \rho} = \frac{\rho_{aria}}{\rho} \approx 4.4 \times 10^{-4}$

$g(0) = \frac{GM}{R_T^2}$

$g(h) < g(0)$

$h \ll R_T$ $\frac{h}{R_T} \ll 1$

$g(h) = \frac{GM}{(R_T+h)^2} = \frac{GM}{R_T^2 \left(1 + \frac{h}{R_T}\right)^2} = \frac{GM}{R_T^2} (1+\epsilon)^{-2} \approx g(0) \left(1 - 2\frac{h}{R_T}\right) = g(0) (1 - 4.7 \times 10^{-6})$

$f(\epsilon) = (1+\epsilon)^{-2} \approx 1 - 2\epsilon$
 $-2(1+\epsilon)^{-3}$

15 m
 6400 km